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Name \_\_\_\_\_

Jan 15, 2016 Friday

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School/Department \_\_\_\_\_

**University Physics**

**Final Examination**

Kuang Yaming Honors School, Nanjing University

$$\Gamma(\nu) = \int_0^{\infty} e^{-t} t^{\nu-1} dt, \quad \Gamma(\nu+1) = \nu\Gamma(\nu), \quad \Gamma(1) = 1, \quad \Gamma(1/2) = \sqrt{\pi}.$$

c	h	e	m <sub>e</sub>	u	N <sub>A</sub>	k <sub>B</sub>
3.0×10 <sup>8</sup> m/s	6.63×10 <sup>-34</sup> J·s 1.24×10 <sup>3</sup> eV·nm/c	1.6×10 <sup>-19</sup> C	9.11×10 <sup>-31</sup> kg 0.511MeV/c <sup>2</sup>	931.494MeV/c <sup>2</sup>	6.02×10 <sup>23</sup> mol <sup>-1</sup>	8.617×10 <sup>-5</sup> eV/K

The quark combinations of some particles: p – uud    n – udd    Λ<sup>0</sup> – uds    Σ<sup>+</sup> – uus

Quark symbols	Charge numbers	Spin	Baryon numbers	Strangeness
u	2/3	1/2	1/3	0
d	-1/3	1/2	1/3	0
s	-1/3	1/2	1/3	-1

Quantum numbers of three quarks:

Select five out of the following six problems.

1.(20 pts.) When neutrons are in thermal equilibrium within a reactor, they satisfy the equipartition theorem.

(a) If the temperature of the reactor is 300 K, what is the average kinetic energy  $\bar{\epsilon}$  of such a thermal neutron?

(b) The mass of neutron is known to be 1.0087 u, What is the corresponding de Broglie wavelength  $\lambda$  ?

Solution: (a) According to the equipartition theorem ( or the law of equipartition) of classical statistical mechanics, the average kinetic energy per degree of freedom is  $1/2 k_B T$ . Here, the translational motion of the neutrons is in the 3 dimensional space and the degree of freedom is 3. Thus the average kinetic energy is

$$\bar{\epsilon} = \frac{3}{2} k_B T = \frac{3}{2} \times 8.617 \times 10^{-5} \text{ eV/K} \times 300\text{K} = 3.88 \times 10^{-2} \text{ eV}. \quad (\bar{\epsilon} = \frac{3}{2} k_B T = \frac{3}{2} \cdot \frac{1}{40} \text{ eV} = \frac{3}{80} \text{ eV})$$

(b) Since  $\bar{\epsilon} \ll m_n c^2$ , it is the non-relativistic case. Thus the corresponding average momentum of the neutrons is

$$\bar{p} = \sqrt{2m_n \bar{\epsilon}} \quad \text{and}$$

$$\lambda = \frac{h}{\bar{p}} = \frac{hc}{c\sqrt{2m_n \bar{\epsilon}}} = \frac{hc}{1240 \text{ nm} \cdot \text{eV} \sqrt{2 \cdot 1.0087 \cdot 3.88 \times 10^{-2} \text{ eV}}} = 0.145 \text{ nm}.$$

(For the relativistic case, using the relations  $E = mc^2 = m_0 c^2 + E_K$  (mass energy=rest (mass) energy+kinetic

energy) and  $E^2 = m_0^2 c^4 + p^2 c^2$ , we obtain

$$p = \frac{1}{c} \sqrt{E^2 - m_0^2 c^4} = \frac{1}{c} \sqrt{(E - m_0 c^2)(E + m_0 c^2)} = \frac{1}{c} \sqrt{E_K (E_K + 2m_0 c^2)}$$

and 
$$\lambda = \frac{h}{p} = \frac{hc}{\sqrt{E_K (E_K + 2m_0 c^2)}}.$$

For the non-relativistic case,  $E_K \ll m_0 c^2$ , we have

$$\lambda = \frac{hc}{\sqrt{E_K(E_K + 2m_0c^2)}} = \frac{hc}{\sqrt{2m_0c^2 E_K \left(1 + \frac{E_K}{2m_0c^2}\right)}} \approx \frac{h}{\sqrt{2mE_K}}.$$

For the extreme-relativistic case,  $E = pc$ , we have  $\lambda = \frac{hc}{E}$ .

2. (20 pts.) After long effort, in 1902, Marie and Pierre Curie succeeded in separating from uranium ore the first substantial quantity of radium, one decigram of pure  $\text{RaCl}_2$  (radium chloride). Actually the radium they obtained was the radioactive isotope  $^{226}\text{Ra}$ , which has a half-life of 1600 a.

- (a) How many nuclei of  $^{226}\text{Ra}$  had they isolated?  
 (b) What is the average life time  $\tau$  of the isotope  $^{226}\text{Ra}$ .  
 (c) What is the decay rate  $R$  of their sample at a late time  $t$ ?  
 (The atomic mass of Cl(chlorine) is 35.5u.)

Solution: (a) The number of isolated  $^{226}\text{Ra}$  nuclei is

$$N_0 = \frac{m}{M(\text{RaCl}_2)} N_A = \frac{0.1\text{g}}{(226 + 2 \times 35.5)\text{g/mol}} \times 6.02 \times 10^{23} \text{ mol}^{-1} = 2.0 \times 10^{20}.$$

(b)  $N = N_0 e^{-\lambda t}$ ,  $\lambda = \frac{\ln 2}{t_{1/2}}$ ,  $-dN = \lambda N_0 e^{-\lambda t} dt$

$$\begin{aligned} \tau &= \frac{\int t(-dN)}{N_0} = \int_0^\infty \lambda t e^{-\lambda t} dt = \frac{1}{\lambda} \int_0^\infty u e^{-u} du = -\frac{1}{\lambda} \int_0^\infty u d(e^{-u}) = -\frac{1}{\lambda} u e^{-u} \Big|_0^\infty + \frac{1}{\lambda} \int_0^\infty e^{-u} du \\ &= -\frac{1}{\lambda} e^{-u} \Big|_0^\infty = \frac{1}{\lambda} = \frac{t_{1/2}}{\ln 2} = 2.04 \times 10^3 \text{ a}. \end{aligned}$$

(c)  $R(t) = -\frac{dN}{dt} = \lambda N_0 e^{-\lambda t} = R_0 e^{-\lambda t}$ ,  $\lambda = \frac{\ln 2}{t_{1/2}} = 0.433 \times 10^{-3} \text{ a}^{-1} = 1.38 \times 10^{-11} \text{ s}^{-1}$ ,

$$R_0 = 2.76 \times 10^9 \text{ s}^{-1} \quad (\text{Bq} \quad \text{Becquerel})$$

3.(20 pts.) An electron of a stationary hydrogen atom passes from the fifth level to the ground state, with a photon emitted in the process.

- (a) Calculate the energy  $E_1$  of the ground state and the energy  $E_5$  of the fifth level, according to the Bohr theory.  
 (b) Considering the recoil of the hydrogen atom, what is the energy  $E_\gamma$  of the emitted photon, to the order of  $(E_5 - E_1)^2 / mc^2$  where  $m = 938.782 \text{ MeV}/c^2$  is the rest mass of the hydrogen atom?  
 (c) To the same order of magnitude, what is the recoil energy  $E_{\text{recoil}}$  of the hydrogen atom?

Solution: (a)  $E_n = -13.6 \frac{1}{n^2} \text{ eV}$ ,  $E_1 = -13.6 \text{ eV}$ ,  $E_5 = -0.544 \text{ eV}$ .

(b) The conservation of momentum gives

$$p_\gamma = p_H. \quad (1)$$

The conservation of energy gives

$$\sqrt{m^2 c^4 + p_H^2 c^2} + E_\gamma = mc^2 + (E_5 - E_1), \quad E_\gamma = p_\gamma c.$$

Substituting  $p_H = p_\gamma$  into above equation, we have

$$\sqrt{m^2 c^4 + E_\gamma^2} + E_\gamma = mc^2 + (E_5 - E_1) \quad \text{or} \quad \sqrt{m^2 c^4 + E_\gamma^2} = mc^2 + (E_5 - E_1) - E_\gamma.$$

To square the both sides of above equation, we have

$$m^2 c^4 + E_\gamma^2 = m^2 c^4 + (E_5 - E_1)^2 + E_\gamma^2 + 2mc^2(E_5 - E_1) - 2mc^2 E_\gamma - 2(E_5 - E_1)E_\gamma$$

or

$$2[mc^2 + (E_5 - E_1)]E_\gamma = (E_5 - E_1)^2 + 2mc^2(E_5 - E_1) - 2(E_5 - E_1)E_\gamma \\ = 2[mc^2 + (E_5 - E_1)](E_5 - E_1) - (E_5 - E_1)^2.$$

Thus

$$E_\gamma = E_5 - E_1 - \frac{(E_5 - E_1)^2}{2[mc^2 + (E_5 - E_1)]} \approx E_5 - E_1 - \frac{(E_5 - E_1)^2}{2mc^2} \quad (\text{Noting that } E_5 - E_1 \ll mc^2)$$

(c)

$$E_{\text{recoil}} = \sqrt{m^2 c^4 + p_H^2 c^2} - mc^2 = \sqrt{m^2 c^2 + p_\gamma^2 c^2} - mc^2 = mc^2 \sqrt{1 + \frac{p_\gamma^2}{m^2 c^2}} - mc^2 \\ \approx \frac{p_\gamma^2}{2m} = \frac{E_\gamma}{2mc^2} \approx \frac{(E_5 - E_1)^2}{2mc^2}.$$

4.(20 pts.) Black body radiation can be treated as an ideal gas of photons. As such, its density of states  $D(\nu)$  in the frequency  $\nu$  space is  $D(\nu) = V 8\pi\nu^2 c^{-3}$  and the number of quantum states within the interval of frequency  $\nu$  to  $\nu + d\nu$  is  $V 8\pi\nu^2 c^{-3} d\nu$ , where  $V$  is the volume of the photon gas.

(a) Show that the internal energy of the photon gas is  $U = \alpha T^4 V$  where  $T$  is temperature of the system and  $\alpha = k_B^4 \Gamma(4) \zeta(4) / \pi^2 c^3 \hbar^3$ .

(b) Find the heat capacity at constant volume  $C_V$  of the system.

(c) Find the Helmholtz free energy  $F$  and the Gibbs free energy  $G$  of the system.

(The fundamental equation of thermodynamics here can be written as  $dU = TdS - pdV$ .)

$$\int_0^{+\infty} \frac{x^{\nu-1}}{e^x - 1} dx = \Gamma(\nu) \zeta(\nu), \nu > 0. \quad )$$

Solution: (a)

$$U = \int_0^\infty \frac{h\nu}{e^{\beta h\nu} - 1} D(\nu) d\nu = \frac{8\pi h V}{c^3} \int_0^\infty \frac{\nu^3}{e^{\beta h\nu} - 1} d\nu = \frac{8\pi h V}{c^3} \frac{1}{(\beta h)^4} \int_0^\infty \frac{x^3}{e^x - 1} dx \\ = \frac{8\pi V k_B^4 T^4}{c^3 h^3} \int_0^\infty \frac{x^3}{e^x - 1} dx = \frac{8\pi V k_B^4 T^4}{c^3 h^3} \Gamma(4) \zeta(4) = \alpha T^4 V.$$

$$(b) \quad C_V = \left( \frac{\partial U}{\partial T} \right)_V = 4\alpha T^3 V.$$

(c) Using the relation  $C_V = T \left( \frac{\partial S}{\partial T} \right)_V = 4\alpha T^3 V$ , we have

$$S = \int_0^T \frac{C_V}{T} dT = S = \int_0^T \frac{4\alpha T^3 V}{T} dT = \frac{4}{3} \alpha T^3 V.$$

Therefore, the Helmholtz free energy is

$$F = U - TS = \alpha T^4 V - \frac{4}{3} \alpha T^4 V = -\frac{1}{3} \alpha T^4 V.$$

Using the relation  $dF = -SdT - pdV$ , we have

$$p = -\left(\frac{\partial F}{\partial V}\right)_T = -\frac{1}{3}\alpha T^3$$

Therefore, the Gibbs free energy is

$$G = U - TS + pV = F + pV = 0.$$

(Or  $G = N\mu = 0$ , where  $\mu$  is the chemical potential of the photon gas (per particle) and is zero.)

5. (20 pts.) A one-dimensional simple harmonic oscillator with mass  $m$  and natural angular frequency  $\omega$  is found to stay in the quantum state of following normalized wave function

$$\psi = c\left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2} e^{-\frac{1}{2}\alpha^2 x^2} + \left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2} \alpha x e^{-\frac{1}{2}\alpha^2 x^2},$$

where  $\alpha = (m\omega/\hbar)^{1/2}$ , and  $c$  is a positive constant to be determined.

(a) Determine the constant  $c$ .

(b) Is the  $\psi$  an eigenfunction of the Hamiltonian of the oscillator?

(c) Find  $\langle E \rangle$ , the average energy of the oscillator.

(d) Find  $\Delta E$ , the uncertainty or fluctuation of the energy of the oscillator.

Solution: (a) From

$$\begin{aligned} 1 &= \int_{-\infty}^{\infty} |\psi|^2 dx = \int_{-\infty}^{\infty} \left[ c\left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2} e^{-\frac{1}{2}\alpha^2 x^2} + \left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2} \alpha x e^{-\frac{1}{2}\alpha^2 x^2} \right]^2 dx \\ &= \int_{-\infty}^{\infty} \left( c^2 \frac{\alpha}{\sqrt{\pi}} e^{-\alpha^2 x^2} + 2c \frac{\alpha^2 x}{\sqrt{\pi}} e^{-\alpha^2 x^2} + \frac{\alpha^3 x^2}{\sqrt{\pi}} e^{-\alpha^2 x^2} \right) dx \\ &= \int_{-\infty}^{\infty} \left( c^2 \frac{\alpha}{\sqrt{\pi}} e^{-\alpha^2 x^2} + \frac{\alpha^3 x^2}{\sqrt{\pi}} e^{-\alpha^2 x^2} \right) dx = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} (c^2 e^{-u^2} + u^2 e^{-u^2}) du \\ &= \frac{c^2}{\sqrt{\pi}} \int_{-\infty}^{\infty} e^{-u^2} du + \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} u^2 e^{-u^2} du = \frac{c^2}{\sqrt{\pi}} \int_0^{\infty} t^{-\frac{1}{2}} e^{-t} dt + \frac{1}{\sqrt{\pi}} \int_0^{\infty} t^{\frac{1}{2}} e^{-t} dt \\ &= \frac{c^2}{\sqrt{\pi}} \Gamma\left(\frac{1}{2}\right) + \frac{1}{\sqrt{\pi}} \Gamma\left(\frac{3}{2}\right) = c^2 + \frac{1}{2}, \end{aligned}$$

we obtain  $c^2 = \frac{1}{2}$  or  $c = \frac{\sqrt{2}}{2}$ .

$$(b) \hat{H} = -\frac{\hbar^2}{2m} \frac{d^2}{dx^2} + \frac{1}{2} m\omega^2 x^2$$

$$\text{Noting that } \frac{d^2}{dx^2} \left( e^{-\frac{1}{2}\alpha^2 x^2} \right) = \frac{d}{dx} \left( -\alpha^2 x e^{-\frac{1}{2}\alpha^2 x^2} \right) = -\alpha^2 e^{-\frac{1}{2}\alpha^2 x^2} + \alpha^4 x^2 e^{-\frac{1}{2}\alpha^2 x^2},$$

and

$$\begin{aligned} \frac{d^2}{dx^2} \left( x e^{-\frac{1}{2}\alpha^2 x^2} \right) &= \frac{d}{dx} \left( e^{-\frac{1}{2}\alpha^2 x^2} - \alpha^2 x^2 e^{-\frac{1}{2}\alpha^2 x^2} \right) = -\alpha^2 x e^{-\frac{1}{2}\alpha^2 x^2} - 2\alpha^2 x e^{-\frac{1}{2}\alpha^2 x^2} + \alpha^4 x^3 e^{-\frac{1}{2}\alpha^2 x^2}, \\ &= -3\alpha^2 x e^{-\frac{1}{2}\alpha^2 x^2} + \alpha^4 x^3 e^{-\frac{1}{2}\alpha^2 x^2} \end{aligned}$$

we have

$$\begin{aligned}\hat{H}e^{-\frac{1}{2}\alpha^2x^2} &= -\frac{\hbar^2}{2m}\frac{d^2}{dx^2}\left(e^{-\frac{1}{2}\alpha^2x^2}\right) + \frac{1}{2}m\omega^2x^2\left(e^{-\frac{1}{2}\alpha^2x^2}\right) = \frac{\alpha^2\hbar^2}{2m}e^{-\frac{1}{2}\alpha^2x^2} - \frac{\hbar^2\alpha^4}{2m}x^2e^{-\frac{1}{2}\alpha^2x^2} + \frac{1}{2}m\omega^2x^2e^{-\frac{1}{2}\alpha^2x^2} \\ &= \frac{m\omega}{\hbar}\frac{\hbar^2}{2m}e^{-\frac{1}{2}\alpha^2x^2} - \frac{\hbar^2\left(\frac{m\omega}{\hbar}\right)^2}{2m}x^2e^{-\frac{1}{2}\alpha^2x^2} + \frac{1}{2}m\omega^2x^2e^{-\frac{1}{2}\alpha^2x^2} = \frac{\hbar\omega}{2}e^{-\frac{1}{2}\alpha^2x^2};\end{aligned}$$

$$\begin{aligned}\hat{H}xe^{-\frac{1}{2}\alpha^2x^2} &= -\frac{\hbar^2}{2m}\frac{d^2}{dx^2}\left(xe^{-\frac{1}{2}\alpha^2x^2}\right) + \frac{1}{2}m\omega^2x^2\left(xe^{-\frac{1}{2}\alpha^2x^2}\right) \\ &= \frac{3\alpha^2\hbar^2}{2m}xe^{-\frac{1}{2}\alpha^2x^2} - \frac{\hbar^2\alpha^4}{2m}x^3e^{-\frac{1}{2}\alpha^2x^2} + \frac{1}{2}m\omega^2x^3e^{-\frac{1}{2}\alpha^2x^2} \\ &= \frac{3\frac{m\omega}{\hbar}\hbar^2}{2m}xe^{-\frac{1}{2}\alpha^2x^2} - \frac{\hbar^2\left(\frac{m\omega}{\hbar}\right)^2}{2m}x^3e^{-\frac{1}{2}\alpha^2x^2} + \frac{1}{2}m\omega^2x^3e^{-\frac{1}{2}\alpha^2x^2} = \frac{3\hbar\omega}{2}xe^{-\frac{1}{2}\alpha^2x^2}.\end{aligned}$$

Therefore

$$\hat{H}\psi = \frac{\hbar\omega}{2}\left(\frac{\alpha}{2\sqrt{\pi}}\right)^{1/2}e^{-\frac{1}{2}\alpha^2x^2} + \frac{3\hbar\omega}{2}\left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2}\alpha xe^{-\frac{1}{2}\alpha^2x^2} \neq E\psi,$$

$\psi$  is not an eigenfunction of  $\hat{H}$ .

(c)

$$\begin{aligned}\langle E \rangle &= \int_{-\infty}^{\infty}\psi^*\hat{H}\psi dx \\ &= \int_{-\infty}^{\infty}\left[\left(\frac{\alpha}{2\sqrt{\pi}}\right)^{1/2}e^{-\frac{1}{2}\alpha^2x^2} + \left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2}\alpha xe^{-\frac{1}{2}\alpha^2x^2}\right]\left[\frac{\hbar\omega}{2}\left(\frac{\alpha}{2\sqrt{\pi}}\right)^{1/2}e^{-\frac{1}{2}\alpha^2x^2} + \frac{3\hbar\omega}{2}\left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2}\alpha xe^{-\frac{1}{2}\alpha^2x^2}\right]dx \\ &= \frac{\hbar\omega}{2}\int_{-\infty}^{\infty}\frac{\alpha}{2\sqrt{\pi}}e^{-\alpha^2x^2}dx + 2\hbar\omega\int_{-\infty}^{\infty}\frac{\alpha^2x}{\sqrt{2\pi}}e^{-\alpha^2x^2}dx + \frac{3\hbar\omega}{2}\int_{-\infty}^{\infty}\frac{\alpha^3x^2}{\sqrt{\pi}}e^{-\alpha^2x^2}dx = \frac{\hbar\omega}{2}\cdot\frac{1}{2} + \frac{3\hbar\omega}{2}\cdot\frac{1}{2} = \hbar\omega.\end{aligned}$$

(d)

$$\hat{H}^2\psi = \hat{H}(\hat{H}\psi) = \left(\frac{\hbar\omega}{2}\right)^2\left(\frac{\alpha}{2\sqrt{\pi}}\right)^{1/2}e^{-\frac{1}{2}\alpha^2x^2} + \left(\frac{3\hbar\omega}{2}\right)^2\left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2}\alpha xe^{-\frac{1}{2}\alpha^2x^2},$$

$$\begin{aligned}\langle E^2 \rangle &= \int_{-\infty}^{\infty}\psi^*\hat{H}^2\psi dx \\ &= \int_{-\infty}^{\infty}\left[\left(\frac{\alpha}{2\sqrt{\pi}}\right)^{1/2}e^{-\frac{1}{2}\alpha^2x^2} + \left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2}\alpha xe^{-\frac{1}{2}\alpha^2x^2}\right]\left[\left(\frac{\hbar\omega}{2}\right)^2\left(\frac{\alpha}{2\sqrt{\pi}}\right)^{1/2}e^{-\frac{1}{2}\alpha^2x^2} + \left(\frac{3\hbar\omega}{2}\right)^2\left(\frac{\alpha}{\sqrt{\pi}}\right)^{1/2}\alpha xe^{-\frac{1}{2}\alpha^2x^2}\right]dx \\ &= \left(\frac{\hbar\omega}{2}\right)^2\int_{-\infty}^{\infty}\frac{\alpha}{2\sqrt{\pi}}e^{-\alpha^2x^2}dx + \left(\frac{3\hbar\omega}{2}\right)^2\int_{-\infty}^{\infty}\frac{\alpha^3x^2}{\sqrt{\pi}}e^{-\alpha^2x^2}dx = \frac{5}{4}(\hbar\omega)^2.\end{aligned}$$

$$\Delta E = \sqrt{\langle E^2 \rangle - \langle E \rangle^2} = \frac{\hbar\omega}{2}.$$

6. (20 pts.) For each of the following reactions state whether or not it is forbidden. If it is forbidden, state a conservation law that is violated.

- (a)  $\Lambda^0 \rightarrow p + \pi^0$   
 (b)  $\bar{p} + p \rightarrow \mu^+ + e^-$   
 (c)  $n \rightarrow p + e^- + \nu_e$   
 (d)  $n + \nu_e \rightarrow p + e^-$   
 (e)  $\Sigma^+ \rightarrow \Lambda^0 + \pi^+$

Answer:

- (a)  $\Lambda^0 \rightarrow p + \pi^0$  Q  
 (b)  $\bar{p} + p \rightarrow \mu^+ + e^-$   $L_e$  and  $L_\mu$   
 (c)  $n \rightarrow p + e^- + \nu_e$   $L_e$   
 (d)  $n + \nu_e \rightarrow p + e^-$  not forbidden weak interaction (neutrino)  
 (e)  $\Sigma^+ \rightarrow \Lambda^0 + \pi^+$  not forbidden strong interaction  
 $uus \rightarrow uds + u\bar{d}$  (creation of  $d \bar{d}$  pair)